



Using the IDIA MeerKAT Pipeline

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Jordan Collier | 20 Feb 2019 | MIGHTEE workshop

Our Systems: Software Stack

Astronomy Tools & Libraries

Research Computing

Scientific Libraries

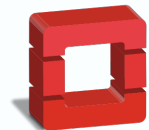
General Purpose Computing



OPEN MPI



Linux



openstack
CLOUD SOFTWARE

ssh access

```
Warning: No xauth data; using fake authentication data for X11 forwarding.
Welcome to Ubuntu 16.04.5 LTS (GNU/Linux 4.4.0-137-generic x86_64)

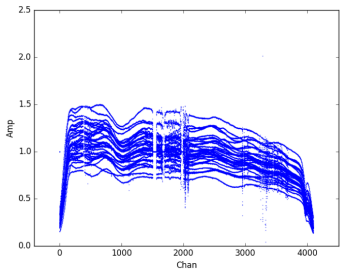
 * Documentation:  https://help.ubuntu.com
 * Management:    https://landscape.canonical.com
 * Support:       https://ubuntu.com/advantage

Get cloud support with Ubuntu Advantage Cloud Guest:
http://www.ubuntu.com/business/services/cloud

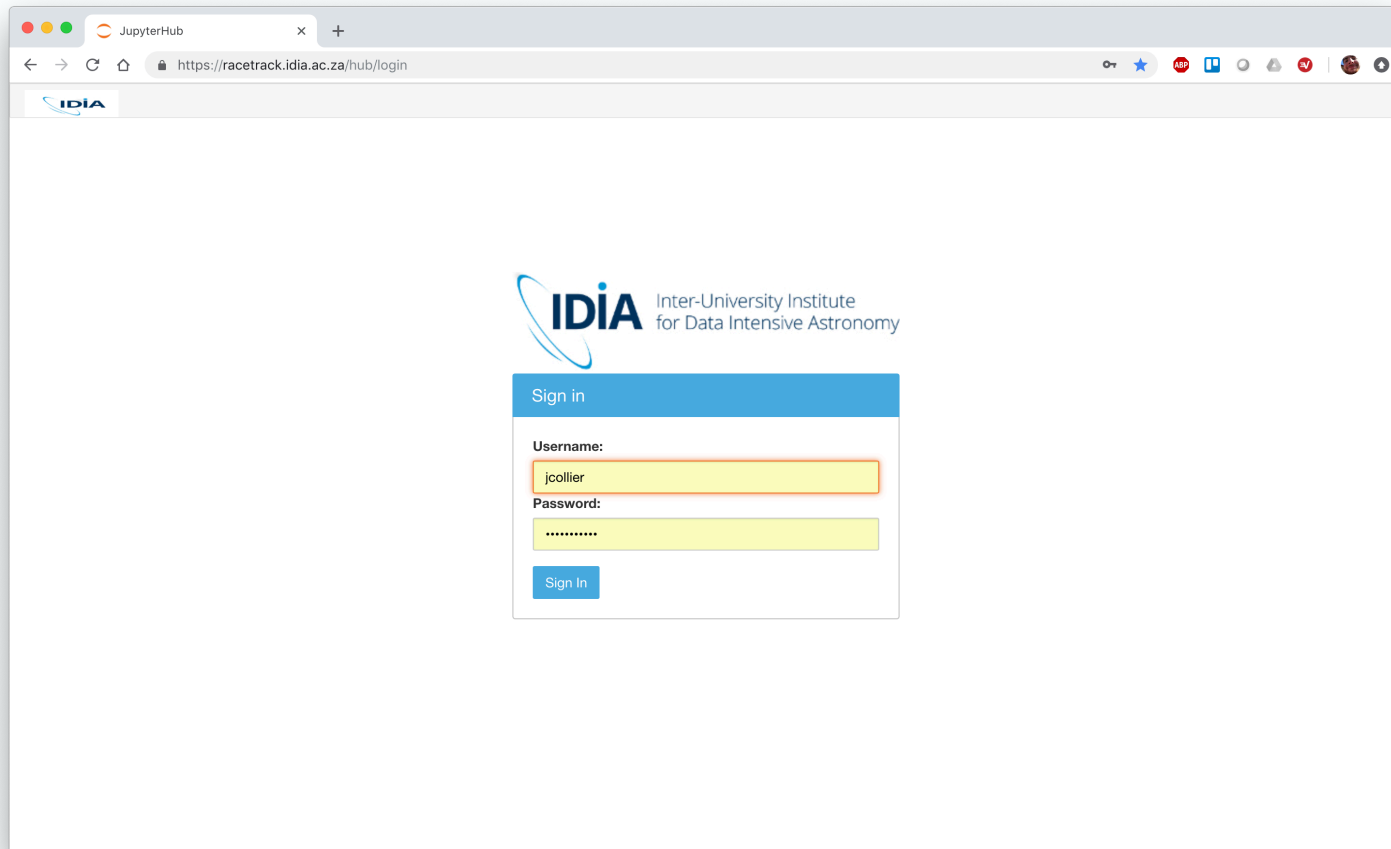
58 packages can be updated.
3 updates are security updates.

New release '18.04.1 LTS' available.
Run 'do-release-upgrade' to upgrade to it.

*** System restart required ***
Last login: Mon Nov 26 02:19:05 2018 from 137.154.142.200
jcollier@racetrack:~$ ls data/CASA_testing/plots/
jcollier@racetrack:~$ ls data/CASA_testing/plot
plotcal.last plots.last plots/
jcollier@racetrack:~$ ls data/CASA_testing/*/plots/
1491550051_bpss_gain.png      1491550051_bpss_real_imag.png      1491550051_fastplot_phase.png      1491550051_gain_phase.png      1491550051_target_spectrum.png
1491550051_bpss_phase.png    1491550051_fastplot_amp.png        1491550051_gain_amp.png          1491550051_target_spec.pdf
1491550051_bpss.png          1491550051_fastplot_phase_benchmarking_helo.png  1491550051_gain_amp.png          1491550051_target_spec.png
jcollier@racetrack:~$ ls data/CASA_testing/*/plots/1491550051_bpss_gain.png
data/CASA_testing/1491550051/plots/1491550051_bpss_gain.png
jcollier@racetrack:~$ imgcat data/CASA_testing/*/plots/1491550051_bpss_gain.png
imgcaty: command not found
jcollier@racetrack:~$ imgcat data/CASA_testing/*/plots/1491550051_bpss_gain.png
```



JupyterHub



JupyterHub

The screenshot displays a JupyterLab environment. On the left, a file browser shows a list of notebooks and files, with 'ms-vis.ipynb' selected. The main area contains a code cell with the following Python code:

```
msvis(MS,col='CPARAM',corrs='all',field='0',xaxis='Real',yaxis='Imag',markersize=1,logy=False,extent=0.1)
msvis(MS,col='CPARAM',corrs='all',field='0',xaxis='Chan',yaxis='Phase',markersize=1,logy=False,extent=0.1)
MS = '/users/jcollier/data/CASA_testing/1491550051/1491550051.g0/'
msvis(MS,col='CPARAM',corrs='all',field='1',xaxis='Time',yaxis='Amp',markersize=1,logy=False,extent=1e-8)
msvis(MS,col='CPARAM',corrs='all',field='1',xaxis='Time',yaxis='Phase',markersize=1,logy=False,extent=1e-8)
#msvis(MS,col='DATA',corrs='all',sortby='ANTENNA1',xaxis='Chan',yaxis='Real',markersize=1)
```

Below the code, the output shows several status messages and data extraction details:

```
Successful readonly open of default-locked table /users/jcollier/data/CASA_testing/1491550051/1491550051.b0/: 13 columns, 16 rows
Extracted data with shape (16, 2, 4096) in 0 seconds.
Plotted 131072 points in 0 seconds.
Wrote figure 'fastplot.png' in 0 seconds.
Successful readonly open of default-locked table /users/jcollier/data/CASA_testing/1491550051/1491550051.b0/: 13 columns, 16 rows
Extracted data with shape (16, 2, 4096) in 0 seconds.
Plotted 131072 points in 0 seconds.
Wrote figure 'fastplot.png' in 0 seconds.
Successful readonly open of default-locked table /users/jcollier/data/CASA_testing/1491550051/1491550051.g0/: 13 columns, 118 4 rows
Extracted data with shape (1, 2, 1184) in 0 seconds.
Plotted 2368 points in 0 seconds.
Wrote figure 'fastplot.png' in 0 seconds.
Successful readonly open of default-locked table /users/jcollier/data/CASA_testing/1491550051/1491550051.g0/: 13 columns, 118 4 rows
Extracted data with shape (1, 2, 1184) in 0 seconds.
Plotted 2368 points in 0 seconds.
Wrote figure 'fastplot.png' in 0 seconds.
```

The bottom part of the output shows a scatter plot with 'Real' on the y-axis (ranging from -1.0 to 2.0) and 'Chan' on the x-axis (ranging from -0.4 to 0.8). The plot contains a dense cloud of blue points centered around the origin.

Singularity Containers

- Led by Jeremy Smith (IDIA, UWC) – slide courtesy
- Containers are a software package that contains everything required to run an application/workflow
 - files, environmental variables, libraries and dependencies
- Allows for interchangeable environments and configurations at runtime and portability between systems



Singularity Containers

```
sourcefinding_py3_update.sh
1 Bootstrap: localimage
2 From: /mnt/containers/sourcefinding_py3_update/sourcefinding_py3-2018-08-30-tmp.simg
3 Include: software-properties-common
4
5 %environment
6
7     export INSTALLER_PATH=/installer
8     ...
9
10 %setup
11 #Copy patch file for blobcat
12 cp blobcat_patch.diff $SINGULARITY_ROOTFS
13
14 %post
15
16 #Installation of BLOBCAT
17 export BLOBCAT_PATH=$INSTALLER_PATH/BLOBCAT
18 mkdir $BLOBCAT_PATH
19 cd $BLOBCAT_PATH
20 wget -O blobcat-v1.2.tar.bz2 http://sourceforge.net/projects/blobcat/files/blobcat-v1.2.tar.bz2/download
21 tar xvjf blobcat-v1.2.tar.bz2
22
23 #Patch blobcat to run python2.7 and fix error on one line
24 patch blobcat.py /blobcat_patch.diff
25
26 ...
27
28 #Upgrade pip
29 pip install --upgrade pip
30
31 #Installation of PASTA
32 pip install git+https://github.com/bwkeller/PASTA
33
34 ...
35
36 #Install eog (and dependencies)
37 apt-get -y install eog
```


Singularity Containers

```
1. IPython: users/jcollier (ssh)
jcollier@racetrack:~$ singularity shell /data/exp_soft/containers/casa-stable-5.3.0.simg
Singularity: Invoking an interactive shell within container...

Singularity casa-stable-5.3.0.simg:~> casa --nologger --nogui --nologfile

=====
The start-up time of CASA may vary
depending on whether the shared libraries
are cached or not.
=====

IPython 5.1.0 -- An enhanced Interactive Python.

CASA 5.3.0-143 -- Common Astronomy Software Applications

--> CrashReporter initialized.
Enter doc('start') for help getting started with CASA...
Using matplotlib backend: TkAgg

CASA <1>: █
```

Singularity Containers

- Exec

```
:~$ singularity exec /data/exp_soft/containers/sourcefinding_py3.simg python myscript.py
```

- MPICASA with SLURM using batch file (“my-casa-job.sh”)

```
#!/bin/bash
#SBATCH --nodes=15
#SBATCH --ntasks-per-node=8
#SBATCH --cpus-per-task=3
#SBATCH --mem=98304
#SBATCH --job-name=tclean
#SBATCH --distribution=plane=4
#SBATCH --output=logs/tclean-%j.out
#SBATCH --error=logs/tclean-%j.err
```

```
/path/to/mpicasa /usr/bin/singularity exec /data/exp_soft/pipelines/casameer-5.4.0.simg "casa"
--nologger --nogui --logfile logs/tclean- $\{\{\text{SLURM\_JOB\_ID}\}\}$ .casa -c tclean.py --config .config.tmp
```

- Terminal on SLURM cluster:

```
user@master001:-> sbatch my-casa-job.sh
```

JupyterHub: Singularity Containers

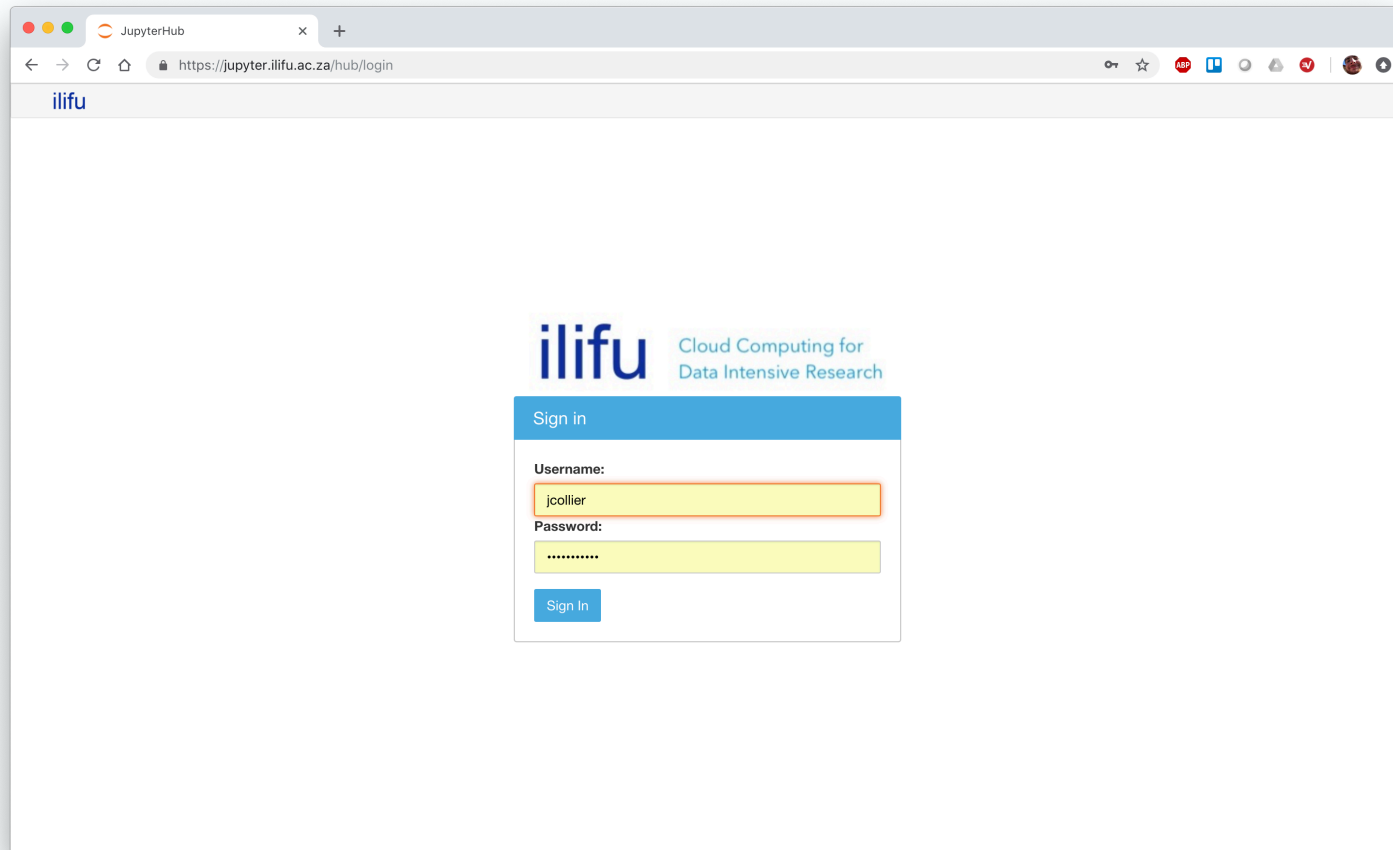
The screenshot displays a JupyterLab interface. On the left, a file browser shows a list of notebooks, with 'ms-vis.ipynb' selected. The main area shows a code editor with Python code for data analysis and plotting. A 'Start Preferred Kernel' menu is open, listing various kernels like 'Jupyter-Casa', 'KERN-2', and 'Py2 Container'. A 'Select Kernel' dialog box is also visible, showing 'Jupyter-Casa' as the selected kernel for the 'ms-vis.ipynb' notebook.

```
[3]: #Works on Kern 2 container
from matplotlib import use
use('TkAgg', warn=False)
import matplotlib.pyplot as plt
from casacore import tables as tb
import numpy as np
import sys
import time
%matplotlib inline

...

[74]: def msvisjez(table, antenna1, antenna2, scan_number, cor_value, adjust_the_image)
    #tb.open(msfile)
    #dat = table.open()
    #Equivalent to (sql-like) command:
    #dat = tb.table('select DATA from $table where ANTENNA1 == $antenna1 and ANTENNA2 == $antenna2 and SCAN_NUMBER == $scan_number and COR_VALUE == $cor_value')
    arr = dat.getcol('CPARAM')
    arr1 = np.real #np.absolute(arr)
    arr1 = arr1.T
    print arr.shape
    fig = plt.figure(figsize=(100, 10))
    #plt.imshow(arr1[cor_value], cmap='gnuplot2', aspect = 0.6)
    #print arr1.shape[1], arr1[0, :].size
    for time in range(arr1.shape[0]):
        for corr in range(arr1.shape[2]):
            plt.scatter(np.arange(0, arr1.shape[1]), arr1[time, :])
            #plt.scatter(np.arange(0, arr1.shape[1]), arr1[time, :])
            #plt.yscale('log')
            #plt.ylim(0, 4.5)
            #plt.gca().invert_yaxis()
    #tb.close()
    plt.savefig('testing.png')
```

JupyterHub Spawner



JupyterHub Spawner

JupyterHub

Home Token Logout

Spawner Options

Select a job profile:

Minimum Node - 1 core, 4 GB, 72 hours

Spawn

- ✓ Minimum Node - 1 core, 4 GB, 72 hours
- Small Node - 2 core, 16 GB, 72 hours
- Medium Node - 4 core, 32 GB, 72 hours
- Large Node - 8 core, 64 GB, 72 hours
- Half-Max Node - 16 core, 125 GB, 72 hours
- Max Node - 32 core, 250 GB, 72 hours
- Max High Memory Node - 32 core, 500 GB, 72 hours



IDiA

Inter-University Institute
for Data Intensive Astronomy

Using the IDIA Pipeline



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processMeerKAT.py

- Builds and submits pipeline job on IDIA / Ilifu SLURM cluster
- Launch pipeline
 - Input measurement set
 - Build your config file
 - OR run your config file (i.e. build job scripts)
 - Request resources
 - Optionally insert your own scripts
 - Specify containers and MPI wrappers

```
1. bash
MC02RTBRAG8WP:~ jordan$ processMeerKAT.py -h
usage: /Users/jordan/Research/MeerKAT/IDIA/pipelines/processMeerKAT/processMeerKAT.py
       [-h] [-M path] [-C path] [-N num] [-t num] [-P num] [-m num] [-p name]
       [-T time] [-S script threadsafe container] [-w path] [-c path] [-l]
       [-s] [-v] (-B | -R | -V)

Process MeerKAT data via CASA measurement set. Version: 1.0

optional arguments:
  -h, --help                show this help message and exit
  -M path, --MS path        Path to measurement set.
  -C path, --config path    Path to config file.
  -N num, --nodes num       Use this number of nodes [default: 8; max: 35].
  -t num, --ntasks-per-node num
                             Use this number of tasks (per node) [default: 4; max:
                             128].
  -P num, --plane num       Distribute tasks of this block size before moving onto
                             next node [default: 2; max: ntasks-per-node].
  -m num, --mem num         Use this many GB of memory (per node) for threadsafe
                             scripts [default: 236; max: 236].
  -p name, --partition name
                             SLURM partition to use [default: 'Main'].
  -T time, --time time      SLURM partition to use [default: 'Main'].
  -S script threadsafe container, --scripts script threadsafe container
                             Run pipeline with these scripts, in this order, using
                             this container (3rd value - empty string to default to
                             [-c --container]). Is it threadsafe (2nd value)?
  -w path, --mpi_wrapper path
                             Use this mpi wrapper when calling threadsafe scripts
                             [default: '/data/exp_soft/pipelines/casa-prerelease-5.3.0-115.el7/bin/mpicasa'].
  -c path, --container path
                             Use this container when calling scripts [default:
                             '/data/exp_soft/pipelines/casameer-5.4.1.xvfb.simg'].
  -l, --local                Build config file locally (i.e. without calling srun)
                             [default: False].
  -s, --submit              Submit jobs immediately to SLURM queue [default:
                             False].
  -v, --verbose              Verbose output? [default: False].
  -B, --build               Build config file using input MS.
  -R, --run                 Run pipeline with input config file.
  -V, --version             Display the version of this pipeline and quit.
```

The Build Step

- `processMeerKAT.py -B -C myconfig.txt -M /data/projects/mightee/MS/1524147354_sdp_10.full.full_pol.MS`
- Builds your config file
 - Reads MS and automatically extracts field IDs based on intent
 - Extracts total flux, bandpass, phase cal, and targets
 - Compares resources you've requested to the number of scans / sub-MSs
 - We partition by scan, so it prevents requesting too many / few threads
 - Ensures reference antenna exists



Config Files

- [crosscal]
 - Select spectral window / freqs (e.g. for HI)
 - Other parameters affecting calibration
- [slurm]
 - Resource parameters
 - List of scripts to run, and whether to call MPI
 - etc
- Path to MS stored
- Field IDs stored

```
myconfig.txt
[crosscal]
minbaselines = 4          # Minimum number of baselines to use while calibrating
specavg = 1              # Number of chans to avg after calibration
timeavg = '8s'          # Time interval to avg after calibration
spw = '0:860~1700MHz'    # spw to use for calibration
calcrefant = True       # Calculate reference antenna in program (overwrites 'refant')
refant = 52             # Reference antenna name/number
standard = 'Perley-Butler 2010' # Flux density standard for setiy
badants = [2, 3, 5, 13, 19, 53] # List of bad antenna numbers in the MS
badfreqranges = [ '944~947MHz', # List of bad frequency ranges
                  '1160~1310MHz',
                  '1476~1611MHz',
                  '1670~1700MHz']

[slurm]
nodes = 8
ntasks_per_node = 2
mem = 236
plane = 1
submit = False
container = '/data/exp_soft/pipelines/casameer-5.4.1.xvfb.simg'
mpi_wrapper = '/data/exp_soft/pipelines/casa-prerelease-5.3.0-115.e17/bin/mpicasa'
partition = 'Main'
time = '12:00:00'
verbose = False
scripts = [('setiy.py', True, ''),
           ('xx_vv_solve.py', False, ''),
           ('xx_vv_apply.py', True, ''),
           ('flag_round_2.py', True, ''),
           ('setiy.py', True, ''),
           ('xy_vx_solve.py', False, ''),
           ('xy_vx_apply.py', True, ''),
           ('split.py', True, ''),
           ('plot_solutions.py', False, ''),
           ('quick_tclean.py', True, '')]

[data]
vis = '/data/projects/mightee/MS/1523541036_sdp_l0.full.full_pol.ms'

[fields]
bpassfield = '0'
fluxfield = '0'
phasecalfield = '1'
targetfields = '2'
```

The Run Step

- `processMeerKAT.py -R -C myconfig.txt`
- Builds SLURM sbatch files (one for each unique python script)
 - Wrote sbatch file "validate_input.sbatch"
 - Wrote sbatch file "partition.sbatch"
 - Wrote sbatch file "calc_refant.sbatch"
 - Wrote sbatch file "flag_round_1.sbatch"
 - Wrote sbatch file "setjy.sbatch"
 - Wrote sbatch file "xx_yy_solve.sbatch"
 - Wrote sbatch file "xx_yy_apply.sbatch"
 - ...
 - Wrote sbatch file "split.sbatch"
 - Wrote sbatch file "plot_solutions.sbatch"
 - Wrote sbatch file "quick_tclean.sbatch"
- Writes master submission (bash) script



Sbatch Files

- Calls script within CASA container, within MPI wrapper
- Logs with job ID written for stdout, stderr, and CASA
- User can choose to submit these jobs one at a time
- OR use master submission script

```
partition.sbatch
#!/bin/bash
#SBATCH --nodes=8
#SBATCH --ntasks-per-node=2
#SBATCH --cpus-per-task=1
#SBATCH --mem=236GB
#SBATCH --job-name=partition
#SBATCH --distribution=plane=1
#SBATCH --output=logs/partition-%j.out
#SBATCH --error=logs/partition-%j.err
#SBATCH --partition=Test02
#SBATCH --time=12:00:00

export OMP_NUM_THREADS=1

/data/exp_soft/pipelines/casa-prerelease-5.3.0-115.el7/bin/
mpicasa singularity exec /data/exp_soft/pipelines/
casameer-5.4.1.xvfb.simg xvfb-run -d casa --nologger! --nogui
--logfile logs/partition- $\{\text{SLURM\_JOB\_ID}\}$ .casa -c
/data/exp_soft/pipelines/dev/processMeerKAT/cal_scripts/
partition.py --config .config.tmp
```

Master Submission Script

- Bash script that submits sbatch files to SLURM queue each time it's run
- Builds ancillary bash scripts that interact with pipeline run
 - Summarise progress
 - Kill all jobs
 - Find errors (after pipeline run)
 - Display runtimes (after pipeline run)

```
submit_pipeline.sh
#!/bin/bash
cp myconfig.txt .config.tmp

#setiy.sbatch
IDS=$(sbatch setiy.sbatch | cut -d ' ' -f4)

#xx_vv_solve.sbatch
IDS+=,$(sbatch -d afterok:$IDS --kill-on-invalid-dep=yes xx_vv_solve.sbatch | cut -d ' ' -f4)

#xx_vv_apply.sbatch
IDS+=,$(sbatch -d afterok:$IDS --kill-on-invalid-dep=yes xx_vv_apply.sbatch | cut -d ' ' -f4)

#flag_round_2.sbatch
IDS+=,$(sbatch -d afterok:$IDS --kill-on-invalid-dep=yes flag_round_2.sbatch | cut -d ' ' -f4)

#setiy.sbatch
IDS+=,$(sbatch -d afterok:$IDS --kill-on-invalid-dep=yes setiy.sbatch | cut -d ' ' -f4)

#xy_vx_solve.sbatch
IDS+=,$(sbatch -d afterok:$IDS --kill-on-invalid-dep=yes xy_vx_solve.sbatch | cut -d ' ' -f4)

#xy_vx_apply.sbatch
IDS+=,$(sbatch -d afterok:$IDS --kill-on-invalid-dep=yes xy_vx_apply.sbatch | cut -d ' ' -f4)

#split.sbatch
IDS+=,$(sbatch -d afterok:$IDS --kill-on-invalid-dep=yes split.sbatch | cut -d ' ' -f4)

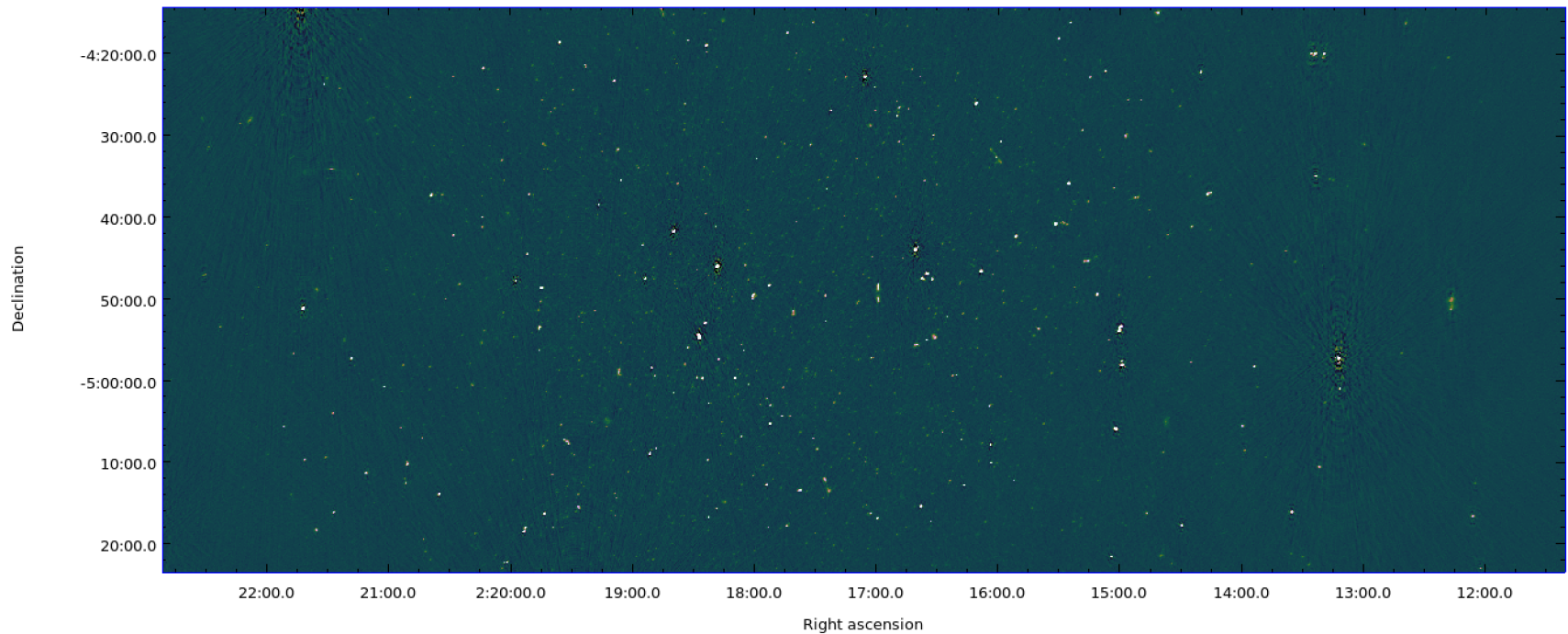
#plot_solutions.sbatch
IDS+=,$(sbatch -d afterok:$IDS --kill-on-invalid-dep=yes plot_solutions.sbatch | cut -d ' ' -f4)

#quick_tclean.sbatch
IDS+=,$(sbatch -d afterok:$IDS --kill-on-invalid-dep=yes quick_tclean.sbatch | cut -d ' ' -f4)

#Output message and create jobScripts directory
echo Submitted sbatch jobs with following IDs: $IDS
mkdir -p jobScripts
```

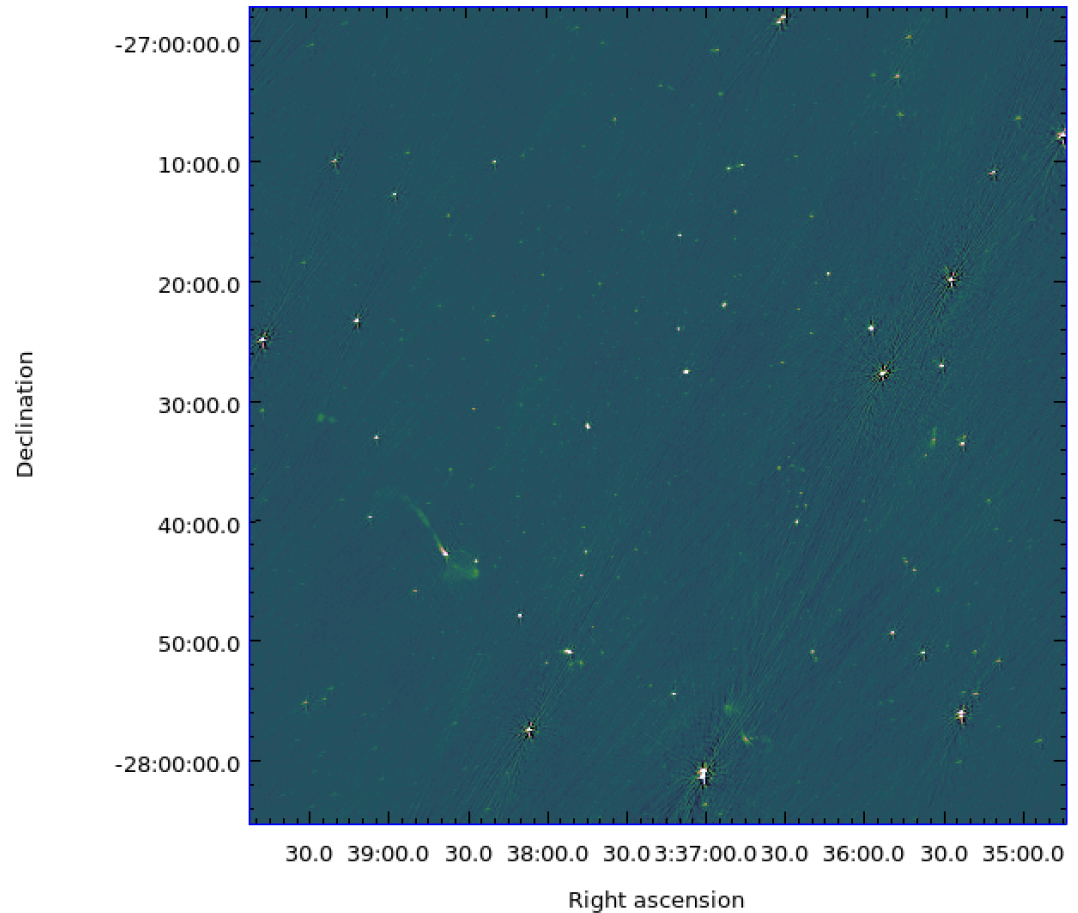
Initial QA: Quick Look Images

- Very quick and dirty imaging for QA purposes
 - No selfcal, no w-projection, no thresholding, no multi-scale, etc
 - XMM-LSS field: RMS ~ 10 μJy / beam



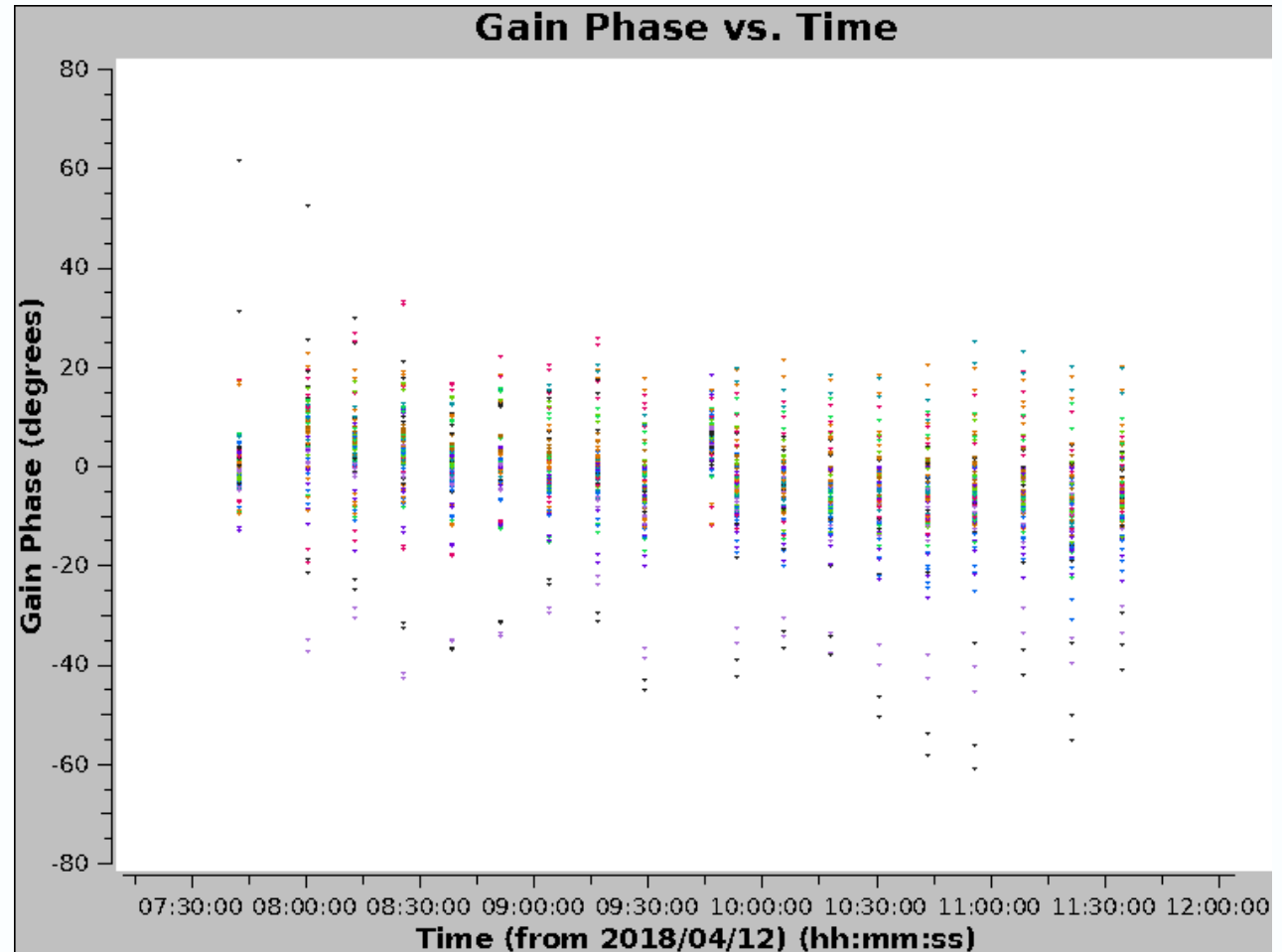
Initial QA: Quick Look Images

- CDFS field
- ~8 hours, 10 MHz spw
- RMS ~80 μ Jy / beam
- Scales to ~10 μ Jy over whole band (assuming ~100 MHz flagged out)



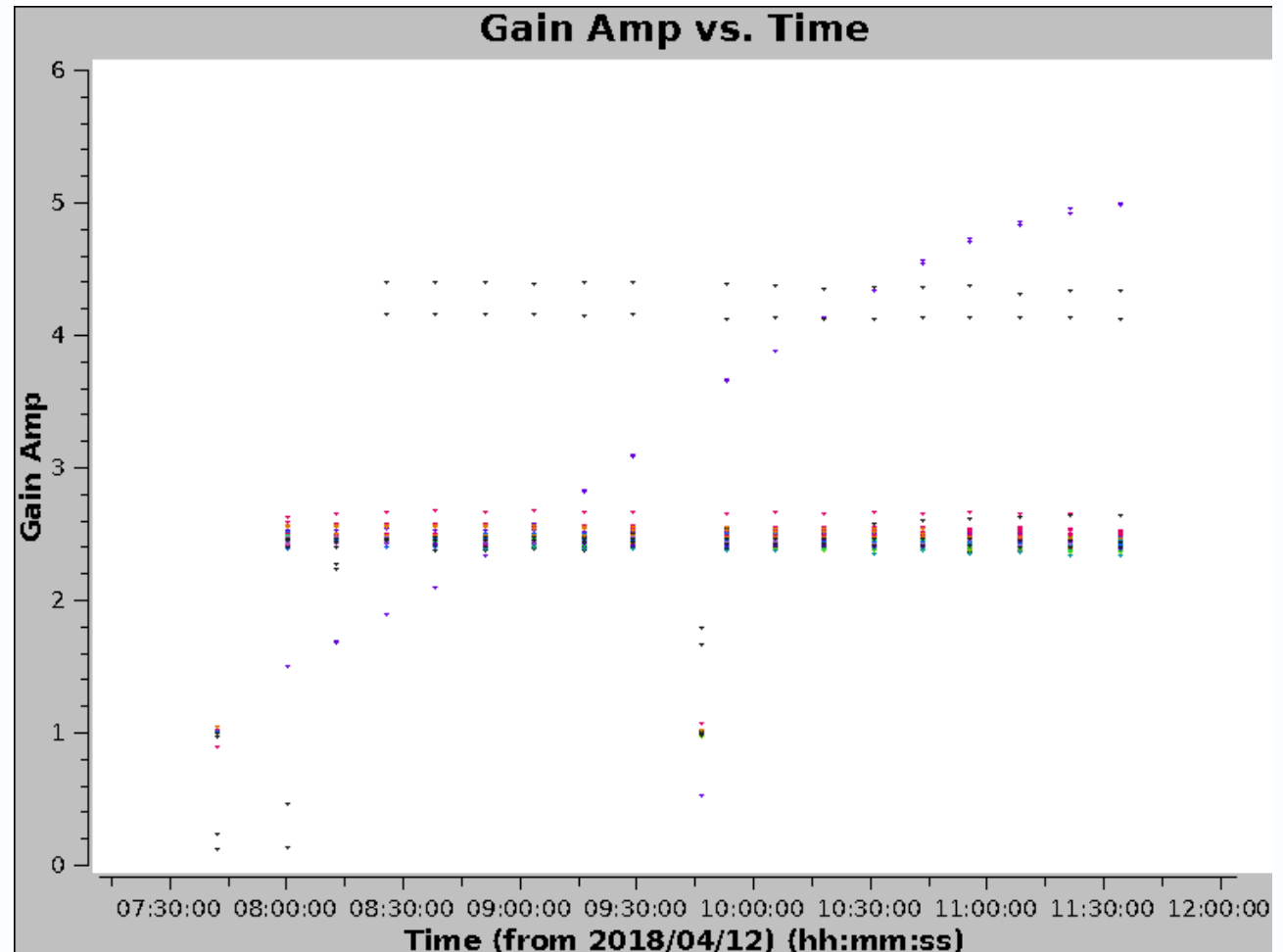
Initial QA: Calibration Solutions

- Phase cal solutions



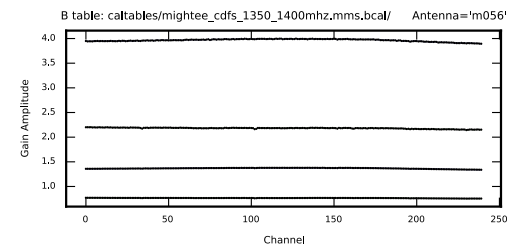
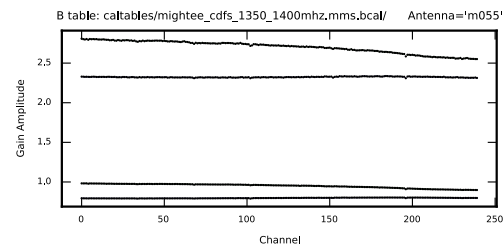
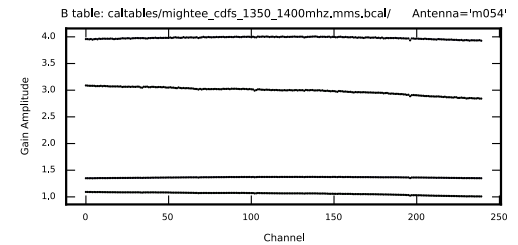
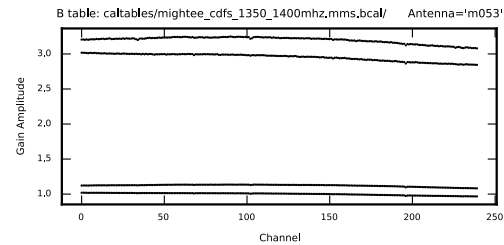
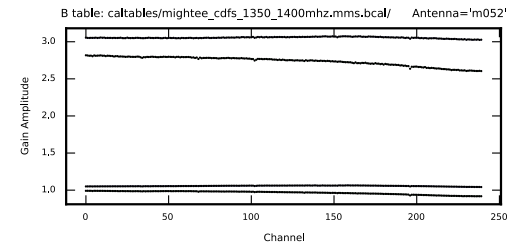
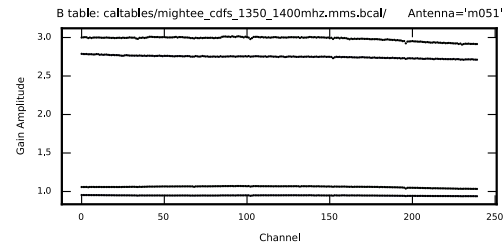
Initial QA: Calibration Solutions

- Phase cal solutions



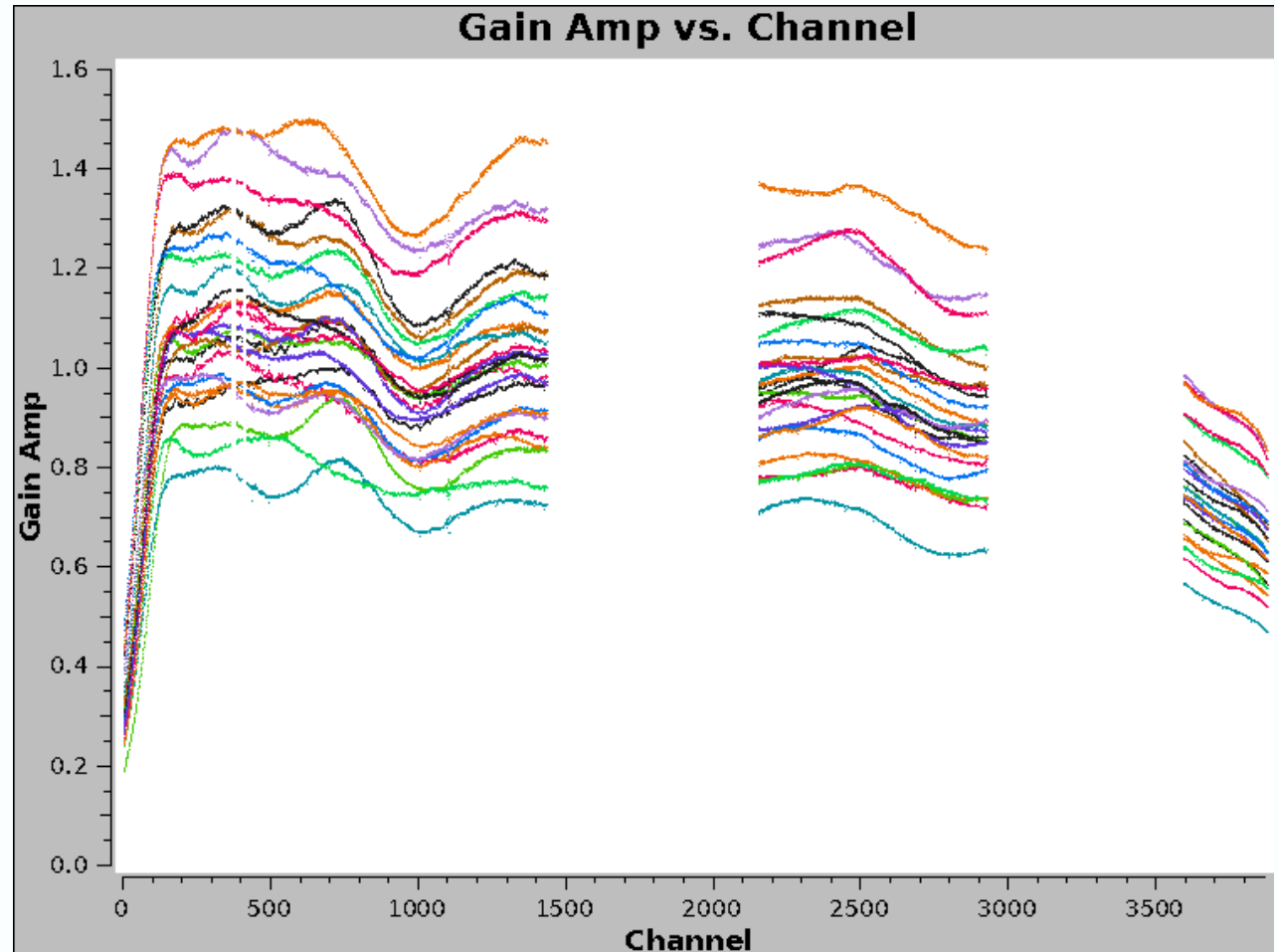
Initial QA: Calibration Solutions

- Phase cal solutions



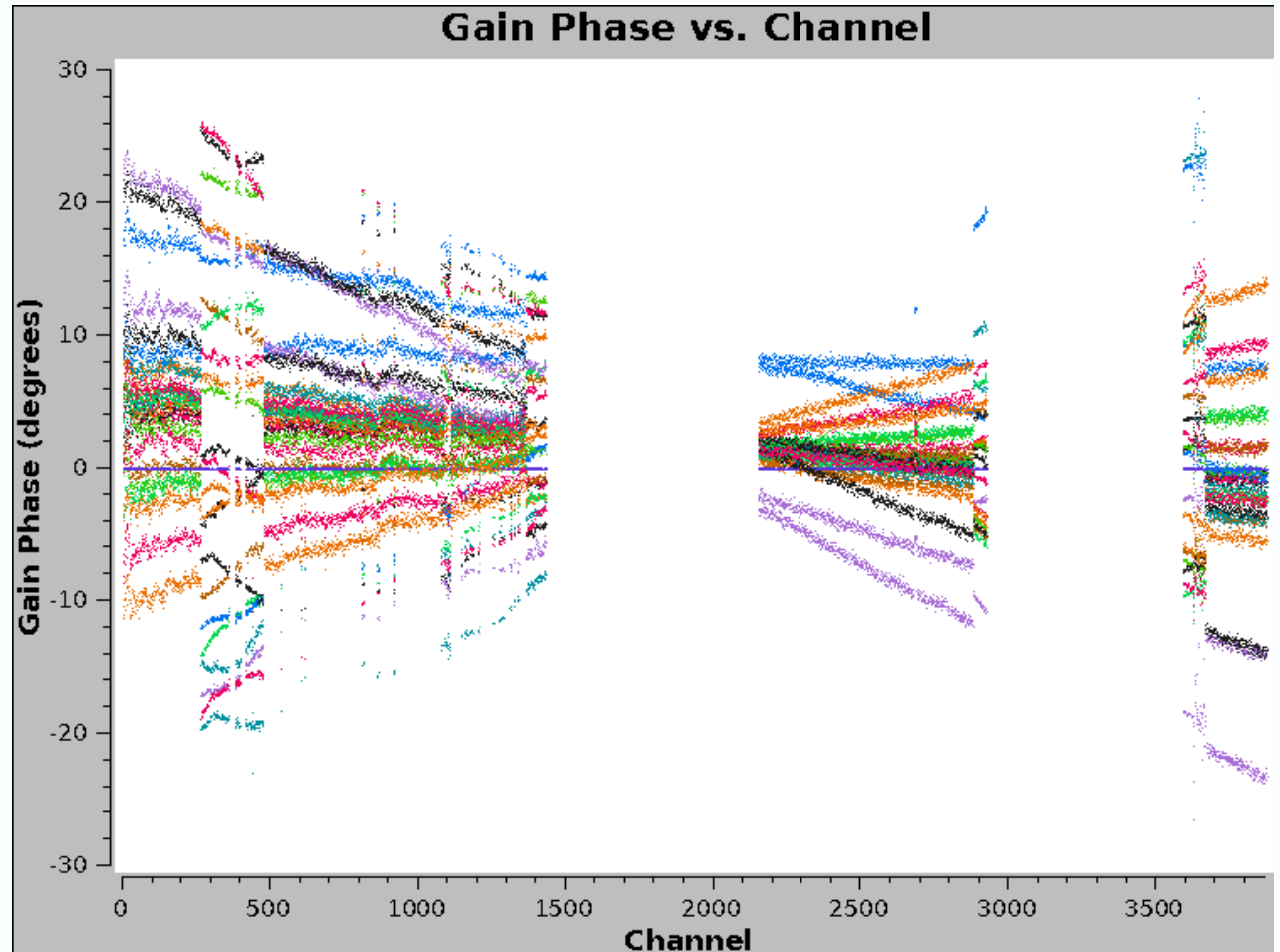
Initial QA: Calibration Solutions

- Bandpass solutions



Initial QA: Calibration Solutions

- Bandpass solutions
- And more...



A Good Framework

- Pipeline outputs calibrated MMS + MSs/MMSs split by field
- You insert your scripts at start, middle or end (e.g. WSClean)
- Each job/script is a logical step that does/doesn't use MPI, and optionally uses a different container
- HPC-friendly – dynamically uses resources & submits to queue
- Use cases we currently support
 - Full stokes calibration + Stokes I only calibration (minimal speedup)
 - Narrow band (spectral line) calibration, full-band calibration
 - Single MS (speedup for small BW), multi-MS
 - Inserting your own scripts (hard-coded or read config file)
- <https://idia-pipelines.github.io/docs/processMeerKAT>

Future Development

- Selfcal and AW Projection (see Krishna's talks)
- Optimisation of resources / performance
 - Currently takes ~1 day to process 64 dish 4k data (~2 TB)
 - Split data by intent during beginning, and simultaneously calibrate / flag
 - More dynamic use of threads and memory per script per intent (based on benchmarking)
 - Partitioning (IO), flagging (RAM), imaging (CPU)
 - Will see a significant speedup, necessary for arrival of 32k data
- Comprehensive data quality assessment



Data Quality Assessment



UNIVERSITY OF CAPE TOWN

IYUNIVESITHI YASEKAPA • UNIVERSITEIT VAN KAAPSTAD

**WESTERN SYDNEY
UNIVERSITY**



Jordan Collier | 20 Feb 2019 | MIGHTEE workshop

Data Quality Assessment

- Already implemented: quality assessment of data products out of processing pipeline
 - i.e. scientific DQA at end of workflow
- End-to-end analysis of MeerKAT (continuum) image / catalogue
- Presents validation report
- Produces several tables summarising data and DQA metrics
- Runs in jupyter notebook on IDIA cloud with interactive plots

MeerKAT Continuum Data Validation Report

Image

File: 'DEEP_2_mfs.sc7.image.tt0.fits'

Date	Field Centre	Central Frequency (MHz)	Synthesised Beam (arcsec)	Median r.m.s. (uJy)	Image peak (Jy)	Dynamic Range (image peak / worst r.m.s.)	Sky Area (deg ²)
2017-04-04T07:48:03.798880:01.1	04:13:26.8223	1273.88	4.4 x 4.3	14	0.03	5E+02	6.27

Catalogue

File: 'DEEP_2_mfs.sc7.image.tt0_aegean_comp.fits'

Source Finder	Flux Type	Number of sources (≥5.0σ)	Multi-component islands	Sum of image flux vs. sum of catalogue flux	Median spectral index	Source Counts Xred ²
aegean	integrated	2769	113	1.1 Jy vs. 1.4 Jy		191.10

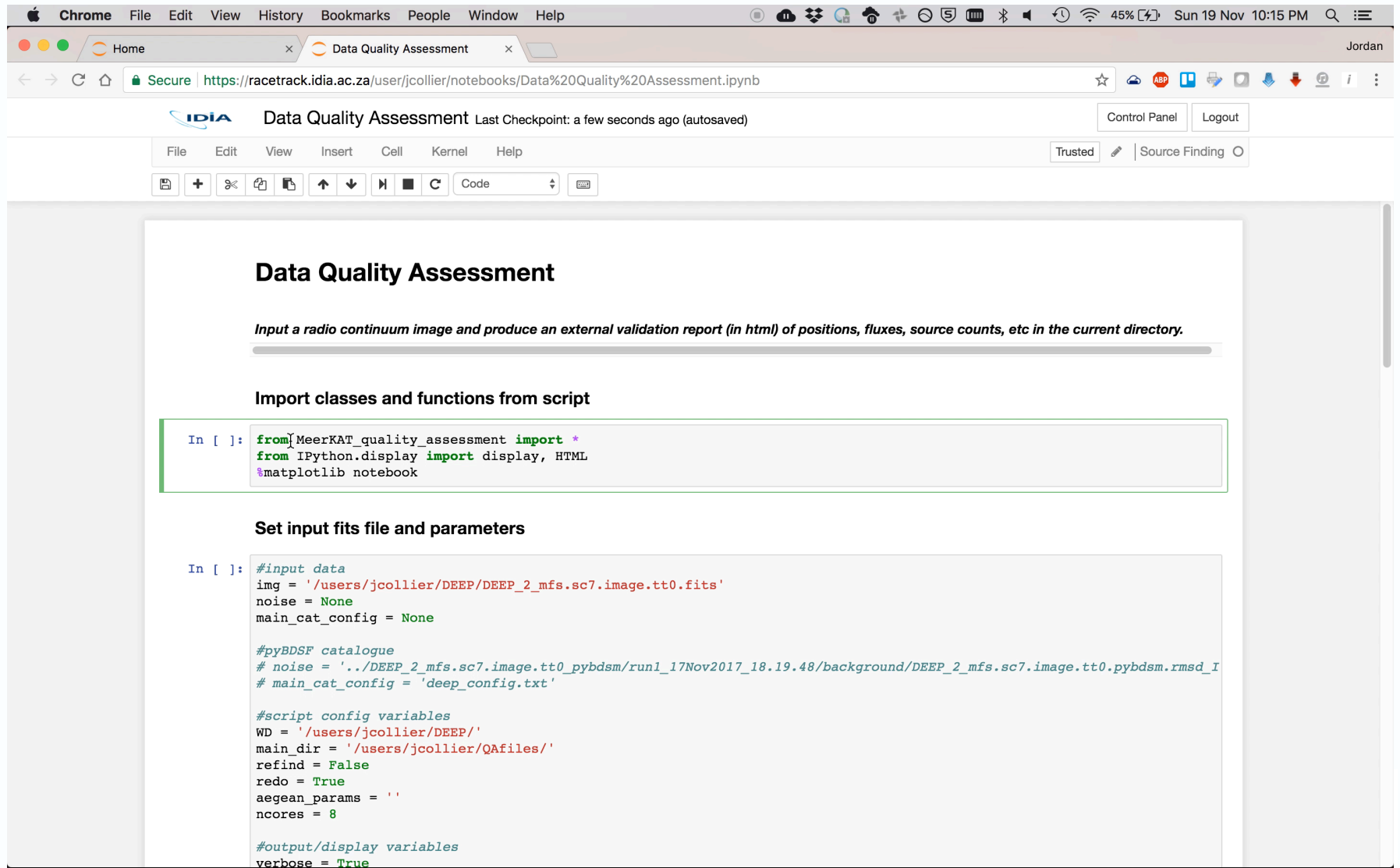
Cross-matches

Survey	Frequency (MHz)	Cross-matches	Median offset (arcsec)	Median flux ratio	Median spectral index
SUMSS	843.0	35	-0.63 ± 2.19 (RA) 0.03 ± 2.85 (Dec)	0.15 ± 0.14 (extrapolated)	-5.40 ± 2.46
GLEAM	201.0	14	-5.72 ± 4.42 (RA) 6.78 ± 10.30 (Dec)	0.04 ± 0.03 (extrapolated)	-2.59 ± 0.52

MeerKAT continuum validation metrics

Flux Ratio (MeerKAT / SUMSS-extrapolated)	Flux Ratio Scatter (MeerKAT / SUMSS-extrapolated)	Positional Offset (arcsec) (MeerKAT - SUMSS)	Positional Offset Scatter (arcsec) (MeerKAT - SUMSS)	Resolved Fraction from int/peak Flux (MeerKAT)	Spectral Index (MeerKAT-SUMSS)	Source Counts Xred ² (MeerKAT)
0.15	0.14	0.63	3.60	0.21	-5.40	191.10

Data Quality Assessment



Chrome File Edit View History Bookmarks People Window Help

Home x Data Quality Assessment x Jordan

Secure | https://racetrack.idia.ac.za/user/jcollier/notebooks/Data%20Quality%20Assessment.ipynb

IDIA Data Quality Assessment Last Checkpoint: a few seconds ago (autosaved) Control Panel Logout

File Edit View Insert Cell Kernel Help Trusted Source Finding

Data Quality Assessment

Input a radio continuum image and produce an external validation report (in html) of positions, fluxes, source counts, etc in the current directory.

Import classes and functions from script

```
In [ ]: from MeerKAT_quality_assessment import *
        from IPython.display import display, HTML
        %matplotlib notebook
```

Set input fits file and parameters

```
In [ ]: #input data
img = '/users/jcollier/DEEP/DEEP_2_mfs.sc7.image.tt0.fits'
noise = None
main_cat_config = None

#pyBDSF catalogue
# noise = '../DEEP_2_mfs.sc7.image.tt0_pybdsm/run1_17Nov2017_18.19.48/background/DEEP_2_mfs.sc7.image.tt0_pybdsm.rmsd_I'
# main_cat_config = 'deep_config.txt'

#script config variables
WD = '/users/jcollier/DEEP/'
main_dir = '/users/jcollier/QAfiles/'
refind = False
redo = True
aegean_params = ''
ncores = 8

#output/display variables
verbose = True
```


Data Quality Assessment

- Have started looking at DQA of early pipeline data products
- In close collaboration with MeerKAT SDP.
 - Framework to measure quality of pipelines.
 - Standard set of metrics between all pipelines
 - Defer processing to most efficient pipelines?
 - Mapping science requirements from LSPs to technical requirements for pipelines

Data Quality Assessment

Table 1: Selected requirements and specifications from various MeerKAT imaging LSPs

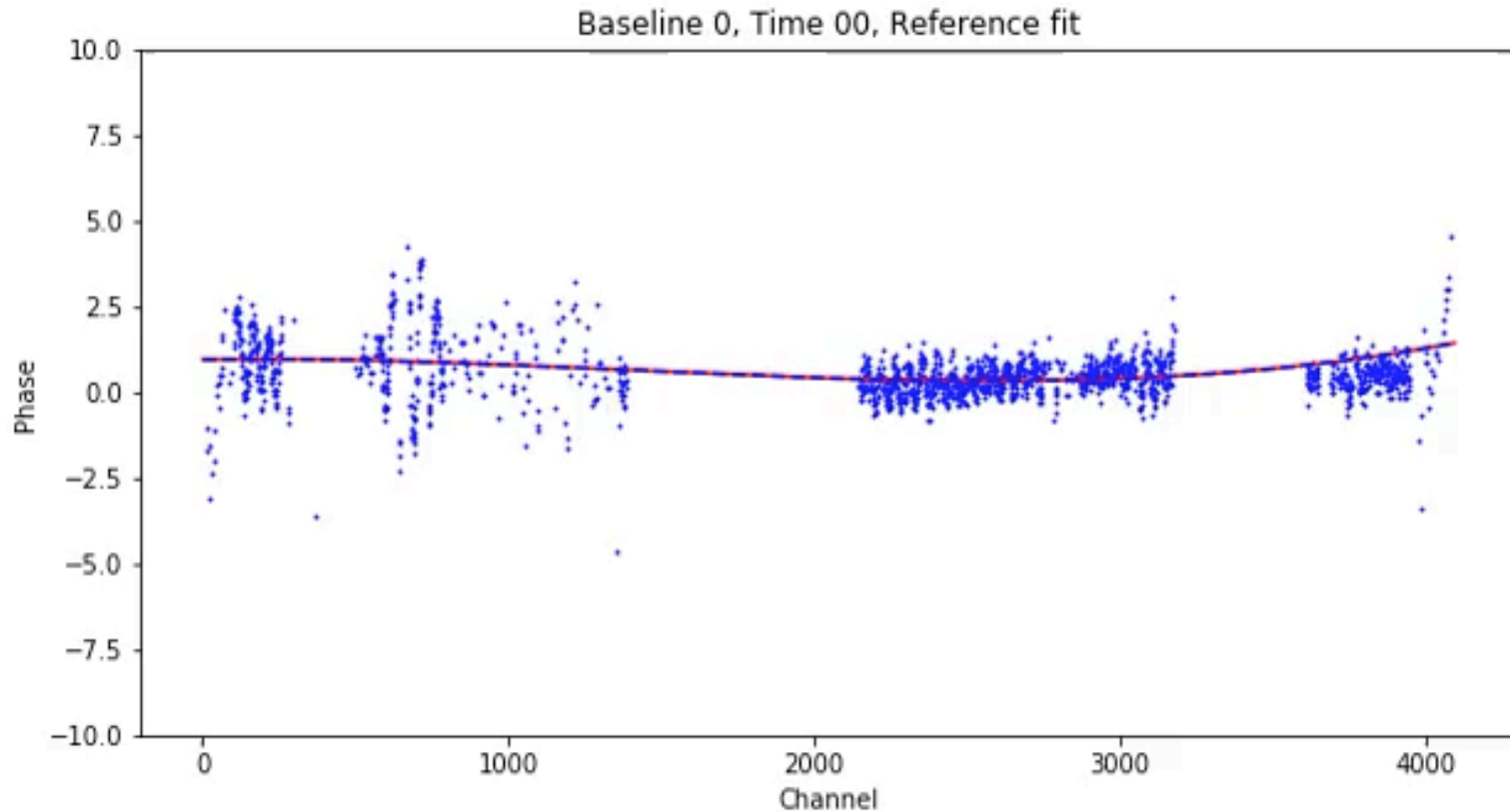
LSP	Sensitivity ($\mu\text{Jy beam}^{-1}$)	Dynamic Range	Velocity Resolution (km s^{-1})	Redshift Range	Area (deg^2)	RM precision (rad m^{-2})	Column Density Sensitivity	H I Mass (M_{\odot})
MIGHTEE (L-band)	$2^a / 90^b$	$\geq 10^{5c}$	6^d	$z \lesssim 0.5$	20	~ 1	$\sim 1 M_{\odot} \text{pc}^{-2e}$	$\gtrsim 2 \times 10^{9f}$
MIGHTEE (S-band)	1^g	$\geq 10^{5c}$	-	$z \lesssim 0.5$	5.5	~ 1	-	-
MIGHTEE (UHF-band)	6^a	$\geq 10^{5c}$	-	$z \lesssim 0.5$	3.5	~ 1	-	-
LADUMA (L-band)	45^b	-	6^d	$0 \leq z \leq 0.58$	0.9–2.2	-	-	$\sim 10^{7.5-10.5}$
LADUMA (UHF-band)	26^h	-	8^i	$0.42 \leq z \leq 1.45$	1.8–5.4	-	-	$\sim 10^{9.2-10.5}$
FORNAX	100^j	-	~ 1	$z \sim 0$	~ 12	-	$\sim 0.1^k - 5^l \times 10^{19} \text{cm}^{-2}$	$\gtrsim 5 \times 10^{5l}$
MHONGOOSE	74^j	-	16	$z \sim 0$	$> 45^m$	$\lesssim 1$	$0.55^n - 7.5^o \times 10^{18} \text{cm}^{-2}$	$\sim 10^{6-11}$
MALS (L-band)	$3 / 500^p$	-	~ 5	$0 < z < 0.85^q$	1000^r	-	$> 10^{19} \text{cm}^{-2}$	$\sim 5 \times 10^4$
MALS (UHF-band)	$3 / 600^p$	-	~ 5	$0.4 < z < 1.87^q$	700^r	-	$> 10^{19} \text{cm}^{-2}$	$\sim 5 \times 10^4$
THUNDERKAT	$\sim 1-1000^s$	-	-	-	-	-	-	-

Data Quality Assessment

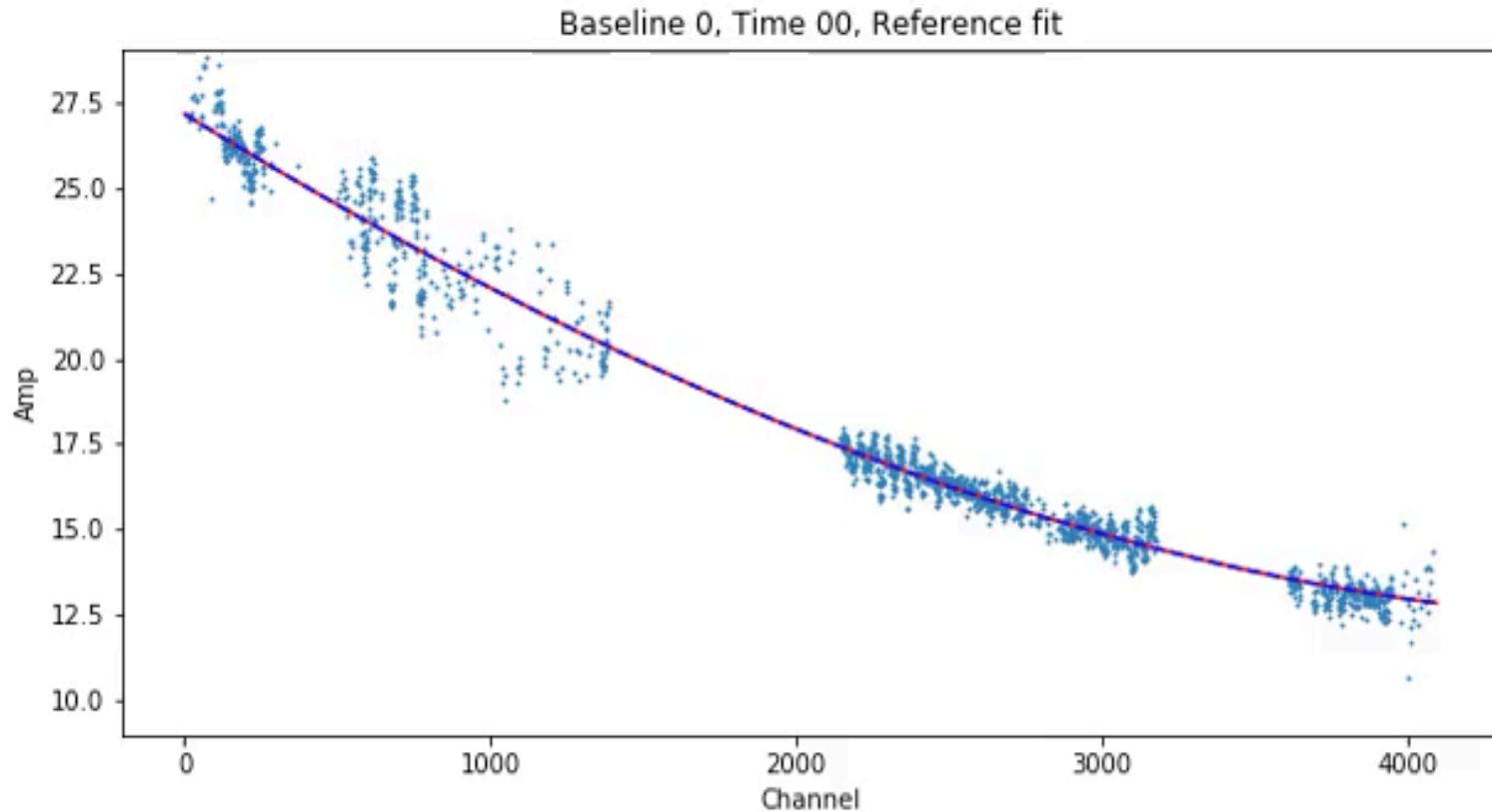
Table 2: Data quality metrics for a simple calibration pipeline. This table is not exhaustive, but represents a selection of metrics we have drafted.

Step	Distribution / plot	Statistic / Metric	Computation	Tolerance
Bandpass calibration	Calibrated bandpass (amplitude and phase) as a function of frequency (per timestamp)	Residual of polynomial fit	Mean reduced χ squared of polynomial fit to each timestamp, compared to reference timestamp ^a	< 10
	Calibrated bandpass (amplitude and phase) as a function of frequency	Normalised median absolute deviation	Maximum normalised absolute deviation from polynomial fit	< 5
Bandpass flagging	Calibrated amplitude as a function of frequency	Fraction of channels flagged	Number of channels with >50% of visibilities flagged divided by total number of channels	< 10%
Phase calibration	Complex gain solutions as a function of time	Outlier metric	Running median?	?
	Calibrated amplitude as a function of frequency	Normalised median absolute deviation	Normalised median absolute deviation from polynomial fit	< 5
Phase cal. flagging	Calibrated amplitude as a function of time	Fraction of timestamps flagged	Number of timestamps with >50% of visibilities flagged divided by total number of timestamps	< 10%
Target flagging	Amplitude as a function of frequency	Fraction of data flagged	Fraction of visibilities flagged divided by total visibilities	< 20%
	Amplitude as a function of frequency	Normalised median absolute deviation	Normalised median absolute deviation	< 5

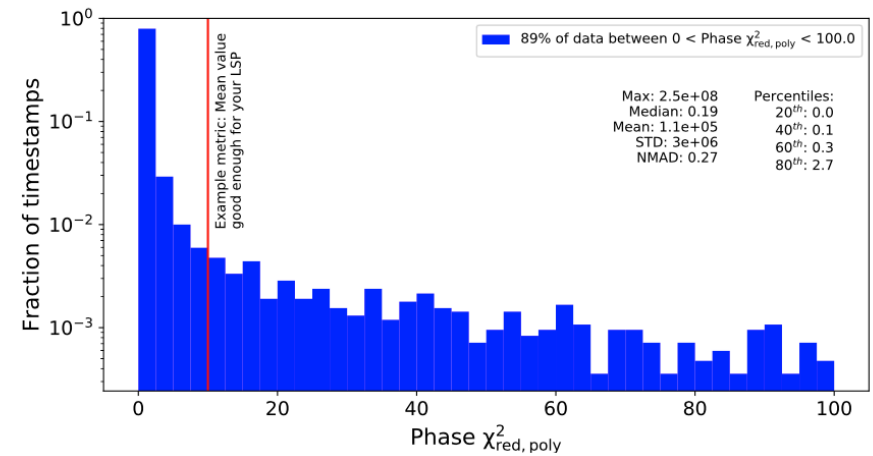
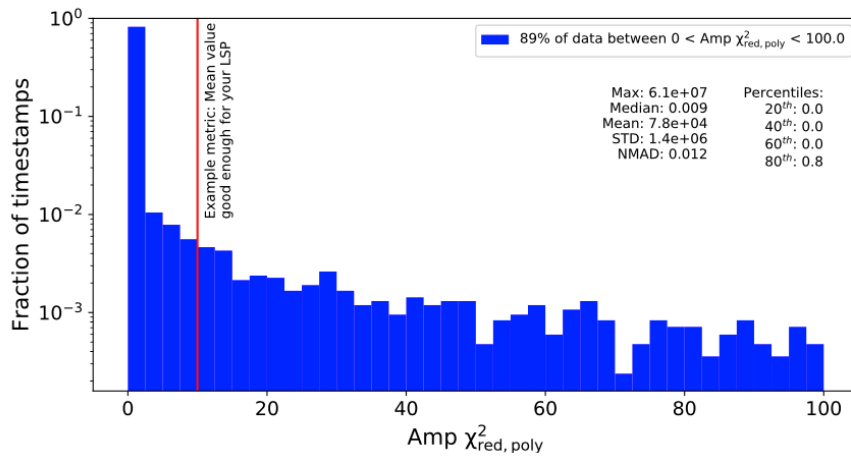
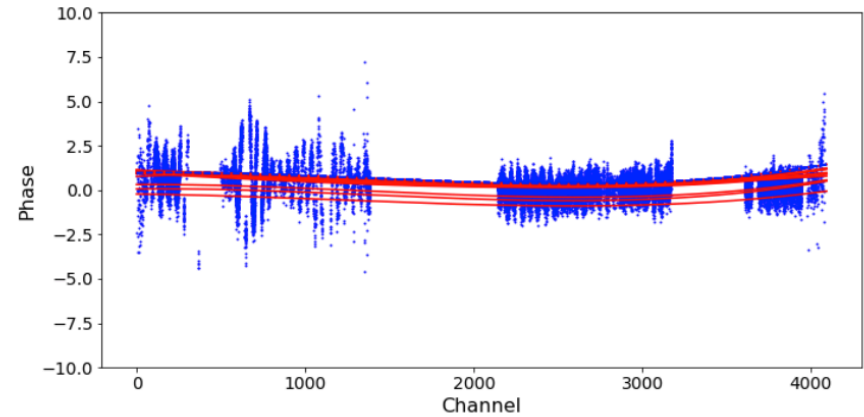
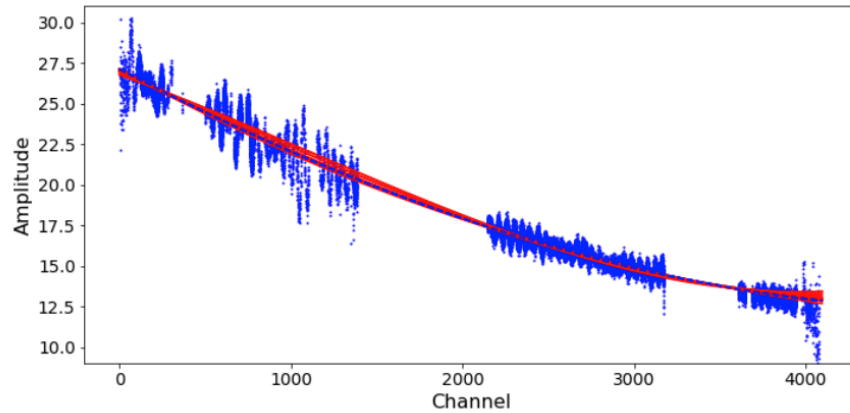
Data Quality Assessment



Data Quality Assessment



Data Quality Assessment



Data Quality Assessment

- DQA being discussed amongst SKA pathfinders in general, within SPARCS DQA WG
 - <http://spacs.pbworks.com/w/page/126067640/dataquality>
- Drafted list of metrics between us
 - https://docs.google.com/spreadsheets/d/1l8x0doPW6LhoWZMLeB5W4tcwzdgw18FWdh5DxnK6_vs/edit#gid=0

Summary

- The IDIA “processMeerKAT” pipeline is an efficient, user-friendly pipeline, that is widely tested and documented
- It runs on the Ilifu cluster, making dynamic use of resources, and containers, and presents a good framework for pipelines
- Many use cases are supported, incl. inserting your own scripts
- Even the quick look images are good
- Coming soon: selfcal, AW projection, optimisation/speedup

THANK YOU

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